

Gravity: How does it Arise from Gauge Theory?

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$$H = tr \left(P_I^2 + [X_I, X_J]^2 + \Psi^* X_I \Gamma' \Psi \right)$$



Matrix Theory Early example of gravity arising from gauge theory: * Large N matrix quantum mechanics * Eigenvalues are positions of particles * Off-diagonal modes are open strings *Quantum fluctuations induce gravity:



't Hooft's argument

* I/N expansion = string perturbation theory

* planar diagram = triangulation of world sheet



Maldacena's argument

)-dimensio

;gs theory

 $(D_v Y)^2 +$

3 of this tl

Extra dimension



Gravitational back reaction ()²+V Y)) creates a warped geometry

$$ds^{2} = a^{2}(r) \eta_{\mu\nu} dx^{\mu} dx^{\nu} + dr^{2} \to e^{-k_{\nu}^{2}} \to e^{-(x_{i}^{\nu} - x_{j}^{\nu})^{2}}$$

Extra dimension = RG scale

Wednesday, May 30, 2012 The (UV regunarized) pranar diagrams of this theory provide a

Large N QCD defines a particular string

theory, in a highly curved target space. In QCD, this string is strongly coupled.





$Gluon = \begin{pmatrix} \bullet & \bullet & \bullet \\ \bullet & \bullet & \bullet \\ \bullet & \bullet & \bullet \end{pmatrix} = Matrix$ $String = (Matrix)^{n}$

QCD String



Large N spin chain = World sheet theory



There are many possibilities:

What is the UV completion of Standard Model? A geometic view: string compact-Standard Model D-brane ification 10¹⁸ GeV

There are many possibilities: String Landscape







- * cut away the AdS space
- * path integral over exterior defines a wave function
- = Wilsonian effective action
 - * radial Schrodinger eqn = Wilsonian RG evolution

$$\hbar \partial_{\Lambda} e^{-\frac{1}{\hbar} S_{grav}(\phi,\Lambda)} = \hat{H}_{grav} e^{-\frac{1}{\hbar} S_{grav}(\phi,\Lambda)}$$



Reconstruct extra dimension from exact RG

Open/Closed String Duality





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UV divergences in a QFT can be thought of as due to on-shell closed strings that propagate in a dual channel



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Counter terms in a (large N) QFT satisfy the eqns of motion of a (classical) closed string field theory.

A schematic derivation, using Polchinski's RG eqn:

$$\delta_{\Lambda} e^{-\frac{1}{\hbar} S_{\text{int}}(A,\Lambda)} = \int \mathcal{D}a \ e^{-\frac{1}{\hbar} \left(\frac{\alpha_{ij}}{2\delta\Lambda} \operatorname{tr}(a^{i}a^{j}) + S_{\text{int}}(A+a;\Lambda)\right)}$$

$$\hbar \partial_{\Lambda} e^{-\frac{1}{\hbar} S_{\rm int}(A,\Lambda)} = \hat{H}_{\rm gauge} e^{-\frac{1}{\hbar} S_{\rm int}(A,\Lambda)}$$

$$\hat{H}_{\text{gauge}} = \hbar^2 \, \alpha^{ij} \, \text{tr} \left(\frac{\partial^2}{\partial A^i \partial A^j} \right)$$

Perform a generalized Hubbard-Stratonovic transform:

$$e^{-\frac{1}{\hbar}S_{\rm int}(A,\Lambda)} = \int \mathcal{D}\phi \ e^{-\frac{1}{\hbar}\left(S_0(A;\phi) + S_{\rm grav}(\phi;\Lambda)\right)}$$

$$S_0(A;\phi) = \sum_{(i_1...i_n)} \phi_{(i_1...i_n)} \operatorname{tr} \left(A^{i_1} \dots A^{i_n} \right)$$

Single trace couplings = closed string fields

Open/closed string duality follows from the identity:

$$\hat{H}_{\text{gauge}} e^{-\frac{1}{\hbar}S_0(A,\phi)} = \check{H}_{\text{grav}} e^{-\frac{1}{\hbar}S_0(A,\phi)}$$

$$\check{H}_{\rm grav} = \beta_I \frac{\partial}{\partial \phi_I} + g_{IJ} \frac{\partial^2}{\partial \phi_I \partial \phi_J}$$

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exact RG evolution = Schrodinger equation

$$\hbar \partial_{\Lambda} e^{-\frac{1}{\hbar} S_{grav}(\phi,\Lambda)} = \hat{H}_{grav} e^{-\frac{1}{\hbar} S_{grav}(\phi,\Lambda)}$$

RG evolution at large N = Hamilton-Jacobi equation

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 * Entanglement entropy is proportional to area: `holography'
 * 5-d Gravity can be understood as a consequence of the thermodynamic laws that relate entropy and energy.

> T. Jacobson E. Verlinde

Jacobson's argument: QFT + Thermo = Einstein Eqn

A) In a Rindler wedge, the vacuum has temperature and entropy:

$$T = \frac{\hbar}{2\pi}, \qquad \qquad S = \alpha A$$

Entropy scales with area because entanglement is dominated by short wavelength modes that cross the Rindler horizon.

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A + B + general geometric facts imply Einstein eqns, with

$$\alpha = \frac{1}{4\hbar G_N}$$

Can we use these ideas to define a quantum gravity in space-time with positive cosmological constant?

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Look for a consistent, covariant cut-off of gauge theory with a holographic dual.

But how do we find it?

Hint from Matrix Theory: BFSS

* Start from a UV complete string theory
* Identify the most basic constituent d.o.f.
* Saturate the number N of constituent d.o.f.
* Take a large N decoupling limit

New idea: Consider U(N N_c) gauge theory on a four sphere Set g = 0 and turn on a maximal homogeneous U(N) instanton flux. So: $U(N N_c) \longrightarrow U(N_c)$



rayior et al (1990)

Yang monopole: homogeneous instanton with maximal charge k New idea: Consider U(N N_c) gauge theory on a four sphere Set $g_{ym} = 0$ and turn on a maximal homogeneous U(N) instanton flux. So: $U(N N_c) \longrightarrow U(N_c)$



LLL dynamics defines a matrix model (ADHM) Take a large N and flat space limit, so that the Planck cells stay finite.

1 ayıdı et al (1990)

Yang monopole: homogeneous instanton with maximal charge k





Given a light-like momentum p and space-time coordinate x

$$p_{AA'} = \tilde{\pi}_A \pi_{A'} \qquad \omega^A = i x^{AA'} \pi_{A'}, \qquad \tilde{\omega}^{A'} = -i \tilde{\pi}_A x^{AA'}.$$

Penrose (1967)

$$\mathbf{PT}_{*} \quad Z^{\alpha} = (\omega^{A}, \pi_{A'}), \qquad \widetilde{Z}_{\beta} = (\widetilde{\pi}_{A}, \widetilde{\omega}^{A'}) \quad \mathbf{PT}^{*}$$

The canonical commutation relations [p,x] = i imply that

Penrose (1967)

$$[Z^{\alpha}, \widetilde{Z}_{\beta}] = \hbar \delta^{\alpha}_{\beta} \qquad \rightarrow \qquad \begin{bmatrix} \omega^{A}, \widetilde{\pi}_{B} \end{bmatrix} = \hbar \delta^{A}_{B} \\ [\pi_{A'}, \widetilde{\omega}^{B'}] = \hbar \delta^{B'}_{A'}$$

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$$[\pi_{A'}, \widetilde{\omega}^{B'}] = \hbar \delta^{B'}_{A'}$$

$$Z^{\dagger}_{\alpha}Z^{\alpha} = \hbar N$$

What does the low energy theory look like?

Lowest Landau Level = Fuzzy Twistor Space!

$$\mathfrak{D} |\Psi\rangle = \hbar N |\Psi\rangle \qquad \qquad \mathfrak{D} = \widetilde{Z}_{\beta} Z^{\beta}$$

This gives a finite Hilbert space of dimension:

$$k = (N+1)(N+2)(N+3)/6.$$

c.f. Zhang & Hu (2002)

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The twistor lines:

$$(\omega^A - ix^{AA'}\pi_{A'})|p\rangle = 0$$

are fuzzy spheres with n = N+1 points

Covariant space-time non-commutativity:

 $,y^{\nu}.$

$$n_i n^i = 1$$

$$\mathbb{CP}^1 \longrightarrow \mathbb{CP}^3 \quad \text{wistor space}$$

$$\int_{\mathbf{V}} S^4 \quad \text{spacetime}$$

$$y^A y_A = \ell^2.$$

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Our Proposal:

UV theory:

N=4 BF Theory with Flux: decoupling limit

IR theory:

ADHM Matrix Model

Emergent theory:

N=4 SYM Theory

Our **Bold** Proposal:

UV theory:

N=4 BF Theory with Flux: decoupling limit

IR theory:

ADHM Matrix Model

Emergent theory:

N=4 SYM Theory

1/N Corrections = Einstein Gravity

Matrix Action = Fuzzy Chern-Simons + Matter

$$\mathcal{A} = \mathcal{A}_I(Z^{\dagger}, Z, \psi^{\dagger}, \psi) d\overline{\mathcal{Z}}^I$$

$$S(\mathcal{A}) = \operatorname{Tr}_{\mathfrak{D} \leq N} \left(\epsilon^{\alpha\beta\gamma\delta} \mathcal{F}_{\alpha\beta} \mathcal{F}_{\gamma\delta} \right)$$

$$\mathcal{F}_{\alpha\beta} = \begin{bmatrix} D_{\alpha}, D_{\beta} \end{bmatrix}, \qquad D_{\alpha} = \frac{1}{\hbar} Z_{\alpha} - \mathcal{A}_{\alpha}$$

$$S_{\text{defect}}(Q, \tilde{Q}, \mathcal{A}) = \text{Tr}\left(\mathcal{I}_{IJ}\widetilde{Q}\mathcal{D}^{I}Q\mathcal{Z}^{J}\right)$$

Why do the 1/N corrections give rise to 4D Einstein gravity?

Hints:

* Jacobson's argument

* Gauge/gravity correspondence

* Twistor string theory contains conformal gravity

* Penrose's `non-linear graviton'

* Gravity MHV amplitudes